

# The Hypothesis of impact origin of diamonds and kimberlites

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## INTRODUCTION

Despite a large volume of accumulated actual material the problem of origin of kimberlites and diamonds still remains unclear. It is reflected in the existence of a number of hypotheses differently and frequently alternatively ascribing the principal aspects of processes of natural kimberlite and diamond formation. Neither has been finally resolved the mechanism of formation of pipes themselves. None of the models offered to this time can adequately explain all features of their structure. The substitution of a cause-and-consequence link, the conditions and mechanisms of formation of diamonds and all types of diamonds and kimberlites is possible in the framework of an impact process – one of the most fundamental geological processes brought about in the Solar system.

## KIMBERLITES, METEORITES, IMPACTITES AND DIAMONDS

**Kimberlite** is a hybrid ultra-basic rock. It is characterized by inequigranular structure caused by xenoliths and megacrysts of refractory minerals (olivine, pyroxene, plagioclase, garnet, picotite) included in the fine-grained groundmass of low temperature minerals (serpentine, carbonate, chlorite). Here the impregnated minerals are represented not by crystals as in conventional magmatic rocks but by rounded "porphyric" grains. And in order to smelt these very refractory minerals, needless to say of the mantle's pressure, the temperatures of more than 2000°C are required. The study of grains' surfaces of these minerals identified a surprising feature: their surface does not practically differ from the river pebble's surface. The so-called "shearless surface" characteristic of the grains of olivine, garnet, ilmenite, apatite as well as the top line and the surface of the Udachnaya pipe, Yakutia – the bottom line, c.f. – details of their surface. [61]

The hybrid nature of kimberlites is expressed by the presence of "related" inclusions in them represented by ultrabasic and eclogitic groups. Their concentration in kimberlites averages about 0.5 vol. % [24]. High-pressure minerals are found in peridotites, dunites and eclogites of phenocrysts. In addition, there is a fine-grained groundmass or in the form of grains and aggregates. The combination of mineral formation processes is characteristic of the inclusions of these rocks to what the zonal crystals of garnet, biotite, olivine testify [17], moreover the crystals of olivine in diamondiferous ultrabasic inclusions are ascribed both with direct and reverse zoning (in respect of Fe/Mg) [38]. The porphyritic structures and the zoning of high-pressure minerals in kimberlites and related inclusions testify to the consistency of rocks in the conditions of fast cooling [17]. Large garnets (1-2 cm) with Cr<sub>2</sub>O<sub>3</sub> content 1-2 mass % and numerous lamellae of clinopyroxene (products of dissolution of solid solution) have been registered in some xenoliths from the Yagerstontep pipe's kimberlites [42]. This testifies to the formation of garnet at the pressure of 10-13 GPa what corresponds to the depth of about 400 km. The autoliths kimberlites (in lamproites similar formations are called lapillite tufts [71]), in which the ellipsoidal or rounded bodies of kimberlite (autoliths) are found in a matrix of fragmented kimberlite and which has a shape of pseudoconglomerate where kimberlite is cemented by kimberlite, is a widely spread variety of kimberlite. Many autoliths have cores around which the accretion of fine-grained kimberlite took place. The cores represent single crystals or small xenoliths of host rocks up to 10 cm in diameter. Likewise formations could have formed only in free space.

The presence in kimberlite of silicate drop-like and dumb-bell-shaped globules (spherules) with an ideally smooth and lustrous surface testifying to their formation in free space is a very interesting phenomenon (Fig. 1,2). Their sizes fluctuate from hundredth parts to 2-3 mm. The color – pale-green, green-to-black. The globules are X-ray-amorphous. Their chemical composition of minerals included in them of CaO and Al<sub>2</sub>O<sub>3</sub>. The composition of gaseous inclusions in the globules is monotonous and rather exotic: N<sub>2</sub> – 57.7-69.9 v.%, hydrocarbons – 20.25 v.%, H<sub>2</sub> – 5 v.%, CO<sub>2</sub> – 10 v.%, [18]. Such globules are known not only in kimberlites. Si carbides, Fe, ferrosilide (FeSi), native Si, Fe, Al are found in spherical exsolutions from diamondiferous tufts of the Northern Urals. A vanadium-bearing globule of silvery-white color is also traced. Its core, contains intermetallic of V<sub>2</sub>Cu<sub>3</sub>Mn<sub>2</sub>Fe<sub>2</sub> composition, the rim is composed of vanadium carbide. The refractoriness of vanadium compounds testify to unusually high temperatures (1900-2800°C) that could be reached at the front of a shock wave. [32]. Similar formations in the form of globules (ore-bearing and silicate) from hundredth parts to several millimeters in size are observed in loose deposits of the known impact craters [26]. They represent the products of evaporation of a meteoroid's body and a target's substance in the impact process and are characteristic of impact structures.

The kimberlite pipes on the surface occupy a small area – from several square meters to 1.5 square kilometers. Three heterogeneous parts – craterial, diametric and bottom are identified within the pipes. The craterial parts of the pipes represent funnels with relatively gently pinching contents filled with disorganized materials. The bottom parts of the pipes are small amount. Their extent to depth reaches several hundreds of meters, the form is generally rounded and oval with rugged profiles. The diametric parts of the pipes have a shape of a pipe with a cylindrical wall with depth. The pipes' walls are basically smooth with longitudinal gliding grooves, generally without signs of contact alterations of pipe-hosting rocks (Fig. 3). The diametric parts of the pipes represent several varieties (intrusion phases) of kimberlite

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"1. The majority of diamonds, if they do not contain a considerable amount of nitrogen, to a certain extent sustained plastic deformations resulting in the series of dislocations in glide planes (111) identified due to double-refraction by an X-ray topographic method using electronic microscope. It means that the diamonds were subjected to an impact of strong displacement stresses at a high temperature (probably more than 2000°C). Other crystals with inequigranular double-refraction are also saturated with dislocations, but in this case the former are polygonized what indicates post-genetic high temperature deformations.

2. The majority of diamonds during the growth sustained numerous, often sharp changes in the chemical composition and medium. Most easily identified evidences to this fact are the layers of growth with wide variations of nitrogen content.

3. In the process of diamond growth two stages of growth and one stage of solution are distinguished: an early stage in the course of which the form of growth was rounded or tubercular-shaped, but poorly developed plane faces of octahedron were also present. In some diamonds this stage is not traced, a stage of growth of diamonds of octahedron habit.

4. The nitrogen is generally present in very thin platelets, meanwhile there are sharp boundaries between the enriched platelets and the localities deprived thereof. It is assumed that the nitrogen must have entered the diamond during its growth at such a temperature at which the diffusion of nitrogen was insignificant and consequently the nitrogen diffused with the formation of platelets. The diffusion must have taken place either in isothermal conditions or in the conditions of considerable temperature increase after the growth; the turbulent growth of diamonds testifies in favor of the second explanation."

The model age of diamond crystals from different regions of the world obtained by a Sm-Nd method owing to subalkaline pyrope-pyroxenes is almost the same and estimated to be 3.1-3.4 billion years (Table 1). These data indicate that the process of global diamond formation was a one-act process not at all associated with the formation of kimberlites themselves.

## Radiological age of diamonds from kimberlites

Table 1

Location of diamond crystals	Age of diamonds (th. years)	Age of kimberlites (mln. years)	Reference
Pipe Kimberli (SAR)	3.4	90	55
Pipe Finsch (SAR)	3.2-3.3	90	55
Pipe Udachnaya (Yakutia, RF)	3.1-3.4	360	54
Arcaress (USA)	3.1	100	47

Note: These data are likely to reflect the age of formation of diamonds with isotopic composition  $\delta^{13}C$  from -2 to -10 ‰.

The isotopic composition of diamonds from kimberlites has a wide range ( $\delta^{13}C$  from +2.7 to -34.4 ‰), however for the overwhelming majority of diamonds it varies within rather small limits ( $\delta^{13}C$  from -1.9 to -9.7 ‰). One should admit that the small range of  $\delta^{13}C$  values is characteristic not only of diamonds from kimberlites but also of diamonds from the alluvial places of Zaire, Venezuela, Brazil, India, Kalimantan and the Urals as well as of diamonds from meteorites. Nevertheless, there are some exceptions. The diamonds from the Ebekeyah alluvial place have  $\delta^{13}C$  values from -11.5 to -22.2 ‰, but for a gray diamond from the Mir pipe the  $\delta^{13}C$  value makes -32.3 ‰. The  $\delta^{13}C$  for Brazilian caroband ( $\pm 8.0$  ‰) is close to the latter value. However the analysis of isotopic composition of carbon in bitumens contained in kimberlite pipes shows that it differs from that of diamonds, carbonates and graphites, its  $\delta^{13}C$  making from -16.4 to -31.4 ‰ [5].

From the information on isotopic geochemistry of diamonds of much interest are the local investigations into isotopic composition of carbon in separate zones of monocrystals growth. In single diamond monocrystals a regular and gradual movement (from the center towards periphery) of  $\delta^{13}C$  content is marked [59]. The only possible mechanism of directed change of isotopic composition for ages is the formation of diamond monocrystals in one continuous process obeyed to the same objective laws of development.

On the whole, the study of isotopic composition of diamonds and sulfides syncretizes their testifies to their polygenesis, belonging to different rocks that have their own history of development [2]. Of special interest are the results of investigations of diamonds from kimberlites from kimberlites of South Africa, lamproites of Australia, as well as in alluvial diamonds of Russia, South-West and Central Africa, Brazil and North America yielded by M. Ozima with his collaborators [51-53]. The values of measured ratios of helium isotopes in diamonds fluctuate from typically radiogenic to those exceeding the planetary value (characteristic of primary helium of Solar system) and being close to the composition of sun's wind.  $^{10}He/^{4}He$  –  $4 \times 10^5$  and even higher [60] what testifies to the presence of a cosmogenic constituent in them [43].

Stone meteorites by composition are close to ultrabasic and basic Earth's rocks. Only ordinary chondrites are composed of high-temperature compounds represented by non-crystallized aggregates, of oxides, xenoliths, crystals and their fragments enclosed in a fine-grained matrix (Fig. 4). Achondrites represent holocrystalline often brecciated rocks. The most spread mineral of meteorites is olivine, and zoning (both direct and reverse) is often characteristic of its crystals [18]. Chromite, spinel, perovskite, diopside, Ca-Al silicate glass, pliesite, troilite, truenite, coenite, pentlandite and others are most often encountered as secondary and accessory minerals.

Montmorillonite, sepiolite, calcite, magnetite, gypsum are present in the matrix in considerable quantities. Carbon in meteorites is represented in the form of composite organic compounds, hydrocarbons, amorphous and thinly dispersed carbon, graphite and its varieties – clonitane. The characteristic feature of this graphite is its chemical purity. The studies of graphite from extraterrestrial bodies by electronic micro-sounding revealed 100% of C [35]. As an associated mineral to meteorite diamonds it is traced in neogenic diamondiferous titanium placers of the Russian platform [15]. Impact effects on diamonds from meteorites, from crushing and brecciation to strong darkening, recrystallization and melting [8]. This can be complemented by the discoveries in meteorites of high-pressure modifications of carbon – diamond, lonsdaleite, choate and other minerals of impact origin, such as moissanite, ringwoodite, majorite. The data of Rb-Sr, Sm-Nd and Pb-Pb methods to determine the age of differentiated meteorites indicate that they were crystallized 4.54 billion years ago. The lower Rb-Sr ratio (internal isochrone) for ages is a result of reworking of the substance of these meteorites not later than 1.5 billion years after their formation [8].

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**Diamonds** are discovered in several iron meteorites and in the majority of ureilites. The ureilites (about 60 finds from 25 g to 11 kg in weight [40]) represent a special variety of meteorite breccia consisting of olivine-cynovite aggregates and a carbonaceous matrix. The olivine, most distributed mineral of ureilites, is uncommon and contains both calcium (0.26-0.42 % CaO) and chromium (0.6-0.9 % Cr<sub>2</sub>O<sub>3</sub>). The high calcium content testifies to a fast cooling process following high temperatures. The high chromium concentration indicates the conditions of formation under which the greater portion of this element was reduced to Cr<sup>2+</sup>, and a special Fe-Mg zoning. The speed of crystals growth during the explosion may reach 50 m/sec [6]. By the conditions of formation the impactites are subdivided into five facies: a base facies, a bottom streams facies, an aerodynamic outburst facies and teetites [31]. The facies of bottom streams contains alluvial breccia of different dimensionality, melted rocks (tagamites) and mixed (suevites). They form a complex system of impact breccias and suevites. The suevites are breccias of impact breccias identified breccia comprising the fragments of polymorphic impact glass of different size and rock fragments consolidated by the same thinly crushed material (Fig. 5). By a number of lithological and petrographic signs: the ratio and composition of vitrolasts, litholasts, grano- and crystalloclasts, as well as by their granulometry, aggregate state and the character of consolidation of the material with host rocks distinguished within astrolites. The accumulations of accreted lapilli representing rounded formations of a concentrically zoned structure 1-1.5 cm in diameter making up to 20-30 % of the rock volume are often traced in suevites. They consist of fragments of minerals of crystalline and sedimentary rocks (70-80 %) and glasses (10-15 %). The suevites are breccias of impact breccias of different dimensionality consolidated by the same lamellae of pelite dimensionality [20]. The bedded bodies of tagamites sometimes change into sub-vertical bodies. The availability of intersecting injection bodies of tagamites is also a characteristic feature. The contacts between tagamites and suevites may be both sharp (though sometimes uneven) and with gradational character. The suevites are breccias of impact breccias, sometimes small apophyses of granites and others are marked. In separate bodies there are peculiar textures where the rocks of one variety are included in another. In some cases large blocks reaching tens and even first hundreds of meters across somewhat "float" in a mass of breccia of lesser dimensionality. The fragmental material contains rock characterizing all known varieties of impact breccias in the surroundings of an impact crater. The complexity of contact relations, availability of reciprocal inclusions and cross sections, "hot" character of the contact interaction testifying to consolidation of these rocks in high temperature conditions are a common feature for the whole diversity of alluvial rocks. On the contrary, the rocks of aerodynamic facies and the teetites differ in structure already after vitrification and essential cooling of the melt that take place in the process of transportation of the target's rock fragments and those of the melt along the ballistic tracks. It is testified to by "cold" contacts of melted and fragmented formations, the aerodynamic sculpture on the surface of glass-coated bombs and lapilli. The concentration of Ni, Co, Cr in impactites is 1.5-3 times as much and Ir – by an order higher than the background concentrations of crater-hosting rocks.

As a result of impacts by meteoroids the terrestrial surface forms craters filled with a poorly sorted and intensively mixed crushed and melted substance of terrestrial rocks, and to a lesser degree, with impactites. More than 200 impact structures up to 335 km have been discovered on earth of late. The fall of large meteoroids is the most important process in the history of the Earth's surface. The impact process is instantaneous, high peak pressure (from 10 to 1000 GPa) and residual temperature (exceeding 1500°C) may occur. During the impact metamorphism there appear high-pressure phases of a number of compounds (coesite, stishovite, diamond, moissanite, majorite, ringwoodite, choate and others), accompanied by the crushing of minerals, destruction of their crystalline lattice, recrystallization and melting of minerals and rocks. The speed of crystals growth during the explosion may reach 50 m/sec [6]. By the conditions of formation the impactites are subdivided into five facies: a base facies, a bottom streams facies, an aerodynamic outburst facies and teetites [31]. The facies of bottom streams contains alluvial breccia of different dimensionality, melted rocks (tagamites) and mixed (suevites). They form a complex system of impact breccias and suevites. The suevites are breccias of impact breccias identified breccia comprising the fragments of polymorphic impact glass of different size and rock fragments consolidated by the same thinly crushed material (Fig. 5). By a number of lithological and petrographic signs: the ratio and composition of vitrolasts, litholasts, grano- and crystalloclasts, as well as by their granulometry, aggregate state and the character of consolidation of the material with host rocks distinguished within astrolites. The accumulations of accreted lapilli representing rounded formations of a concentrically zoned structure 1-1.5 cm in diameter making up to 20-30 % of the rock volume are often traced in suevites. They consist of fragments of minerals of crystalline and sedimentary rocks (70-80 %) and glasses (10-15 %). The suevites are breccias of impact breccias of different dimensionality consolidated by the same lamellae of pelite dimensionality [20]. The bedded bodies of tagamites sometimes change into sub-vertical bodies. The availability of intersecting injection bodies of tagamites is also a characteristic feature. The contacts between tagamites and suevites may be both sharp (though sometimes uneven) and with gradational character. The suevites are breccias of impact breccias, sometimes small apophyses of granites and others are marked. In separate bodies there are peculiar textures where the rocks of one variety are included in another. In some cases large blocks reaching tens and even first hundreds of meters across somewhat "float" in a mass of breccia of lesser dimensionality. The fragmental material contains rock characterizing all known varieties of impact breccias in the surroundings of an impact crater. The complexity of contact relations, availability of reciprocal inclusions and cross sections, "hot" character of the contact interaction testifying to consolidation of these rocks in high temperature conditions are a common feature for the whole diversity of alluvial rocks. On the contrary, the rocks of aerodynamic facies and the teetites differ in structure already after vitrification and essential cooling of the melt that take place in the process of transportation of the target's rock fragments and those of the melt along the ballistic tracks. It is testified to by "cold" contacts of melted and fragmented formations, the aerodynamic sculpture on the surface of glass-coated bombs and lapilli. The concentration of Ni, Co, Cr in impactites is 1.5-3 times as much and Ir – by an order higher than the background concentrations of crater-hosting rocks.

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